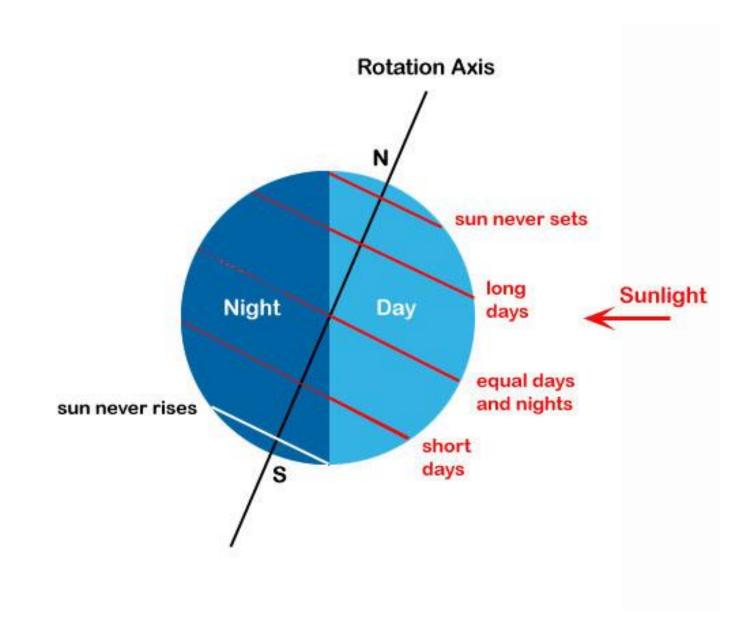
Do you understand the reason for the "seasons"?

Your friend wonders "Why are temperatures on Earth warmer in summer than winter?"

You explain that the tilt of the Earth's spin axis ...

A. causes the Sun to be closer to Earth in summer.

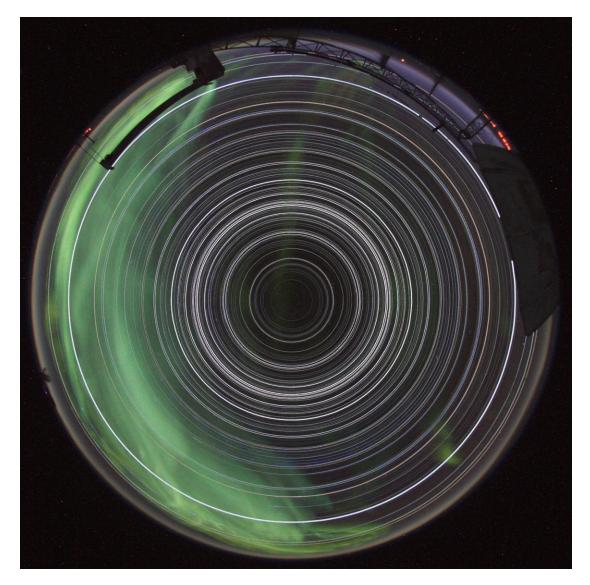
- B. causes sunlight to be concentrated more on the Earth's surface in summer.
- C. causes shorter days in summer.
- D. all of the above
- E. B and C





- Remember to turn in your Doodle sheets at the end of each class!
- *"Aberration of Starlight"* in textbook:
 - Section 2.6.2 (pp. 57-58)
- Install "Stellarium"
- Astronomy Club (4 pm), N210

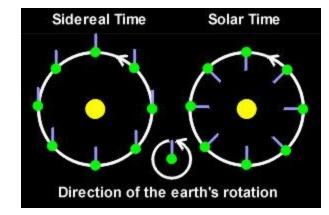
Where was this picture taken? exposure time = ?



What is a "day"? Polya: Understanding the problem Draw a picture.

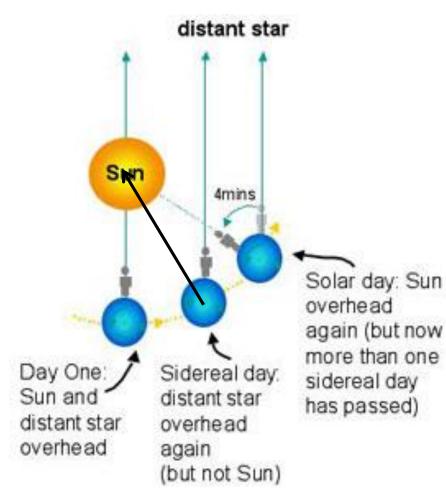
"Understanding the Problem" What is "rotation"? What is a "day"? solar vs. sidereal day

- Period with respect to the Sun
 - 24 hours
 - varies (-18 sec; +30 sec) Why?
- Period with respect to stars ("sidereal")
 - 23^h 56^m 4.1^s
 - 23:56:04.1
 - 23.93 hours



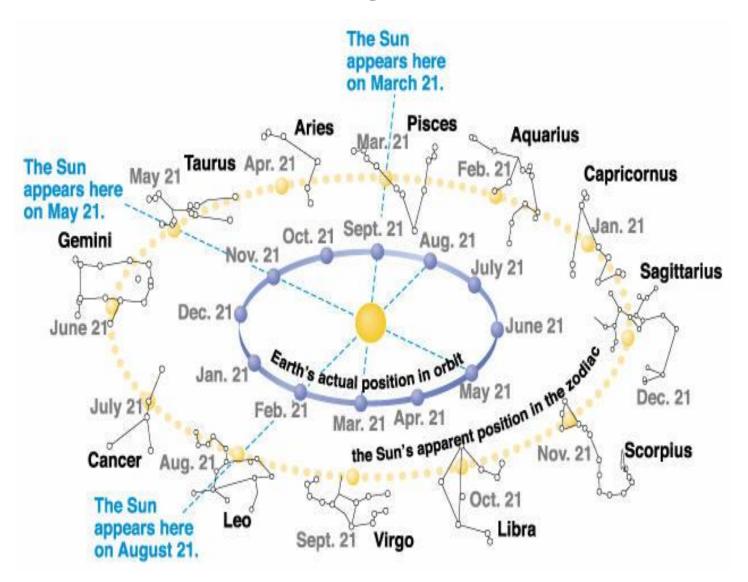


What is a "day"? Draw a figure. assumptions?

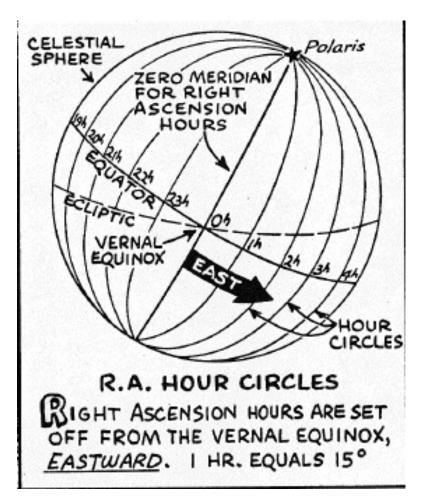


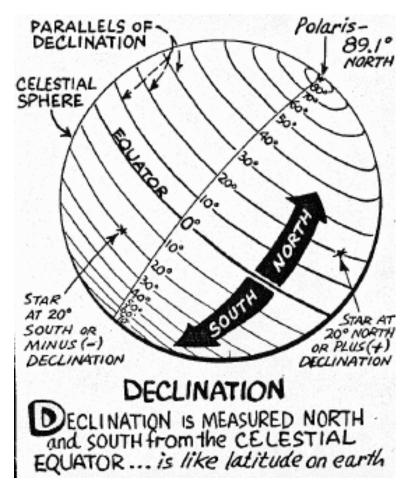
- Earth revolves around the Sun.
 - 360 degrees in 365 days
 - 360 degrees in ~360 days = ~ 1 deg/day
 - to what accuracy?
- What does this mean?
 - Earth must rotate more to "reacquire" the Sun.
 - How long does the Earth require to rotate 1 deg?
 - 360 degrees in 24 hours = ~15 deg/hour
 - How many degrees per minute?
- What information needed for Jupiter?
 - Physical intuition: What do you expect?

"Sidereal Time" at Midnight varies throughout the year



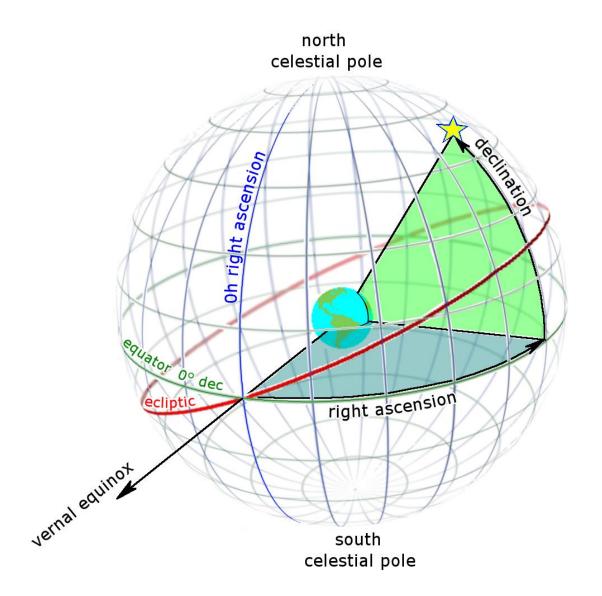
Equatorial Coordinates Right Ascension, Declination





Add "Celestial Equator" as a path on your plastic ball.

Equatorial Coordinates

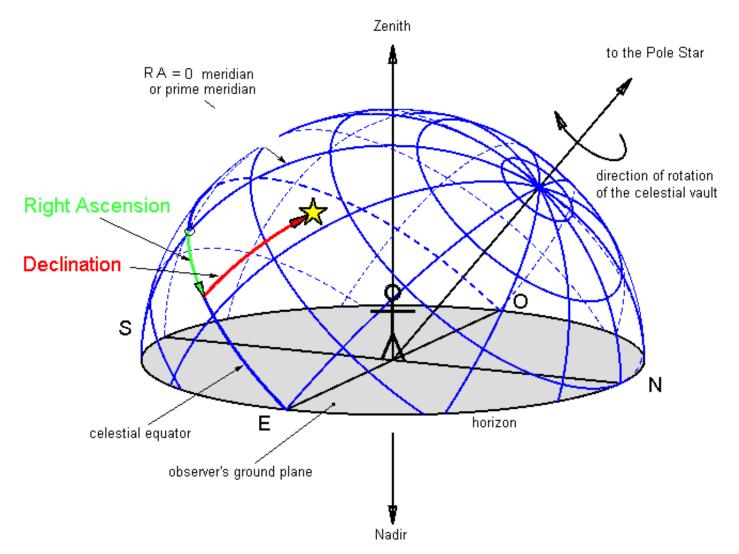


Problem #1 What information do you need?

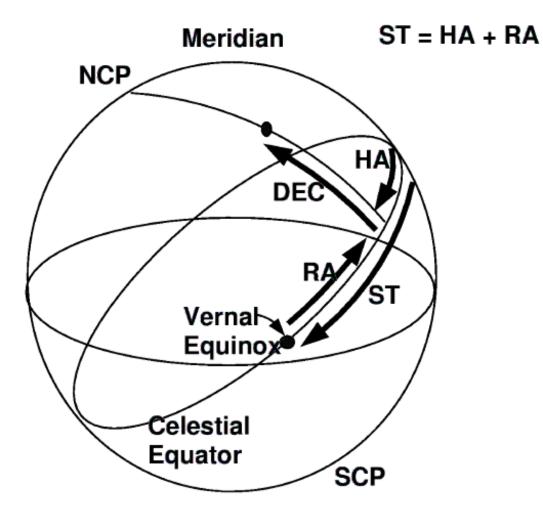
If a Near-Earth Object (aka, NEO) were discovered at midnight tonight, with the following coordinates, from Mt. Lemmon near Tucson, could it be observed simultaneously at Mauna Kea Observatory in Hawaii?

- RA: 22:30:00
- DEC: +20:00:00
- Mauna Kea:
 - Longitude (155.47 deg W)
 - Latitude (19.82 deg)

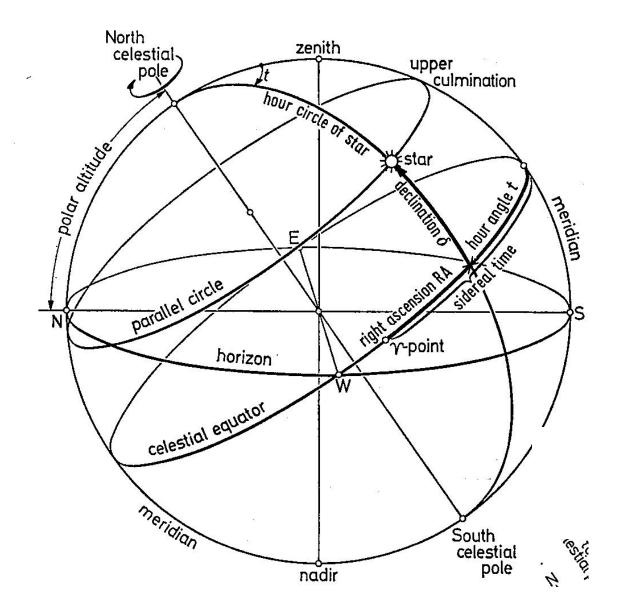
More Terms: sidereal time, hour angle, transit



"Sidereal Time" and "Hour Angle"



"Sidereal Time" and "Hour Angle"



Problem #2 What information do you need?

When is the Galactic Center best studied throughout the night?

- RA: 17:45:40.04 (h: m: s)
- DEC: -29:00:28.1 (deg: arcmin: arcsec)

– LST = 0 on September 26 at midnight

Problem #3 today's optional homework problem

On 17 August 2017, 12:41:04 UT (i.e., Universal Time), the gravitational-wave observatory, LIGO, detected an event (aka, GW170817) from a pair of merging neutron stars. The estimated location had coordinates of (RA: 13:09:48.08; DEC: -23:22:53.3). Quickly, an electronic alert message was sent around the world, encouraging astronomers at observatories around the world to record the event across the spectrum of light. This event provided an opportunity for "multi-messenger" astrophysics.

In Greenwich, the Local Sidereal Time (LST) was 10.35 hours. The 1meter Swope Telescope in Chile (longitude: 70:42:05.9 W; latitude: -29:00:35.85) was the first facility to observe the object. At the time of the event, what was the object's Hour Angle at the Swope Observatory; in other words, could Swope astronomers have observed the event immediately or did they need to wait some amount of time for the object to rise in their sky? [NOTE that Right Ascension increases towards the east.]

Do you understand the problem? What do we know?

When:17 August 2017 at 12:41:04 UTWhere:RA: 13:09:48.08; DEC: -23:22:53.3

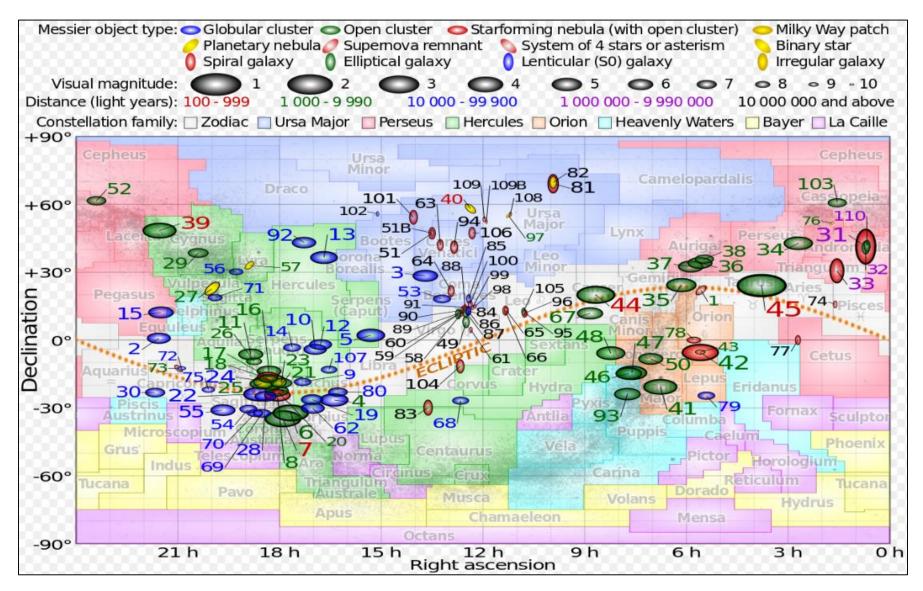
In Greenwich, the Local Sidereal Time (LST) was 10.35 hours.

The 1-meter Swope Telescope in Chile: longitude: 70:42:05.9 W latitude: -29:00:35.85

At the time of the event, what was the object's Hour Angle at the Swope Observatory.

- Longitude: 70:42:05.9 W = 70.70 deg
- 70.70 deg / 15 deg per hour = 4.71 hours
- 4.71 hours earlier than Greenwich
- LST at Greenwich was 10.35 hours at midnight
- At Swope, LST then was 10.35 4.71 = 6.54 hours
- So an RA of 6:32:24 was transiting
- Object's RA = 13:09:48.08
- So HA (east) = 6:37:24.08 = 6.62 hours
- So object had not yet risen above Chilean horizon

"Messier Marathon"

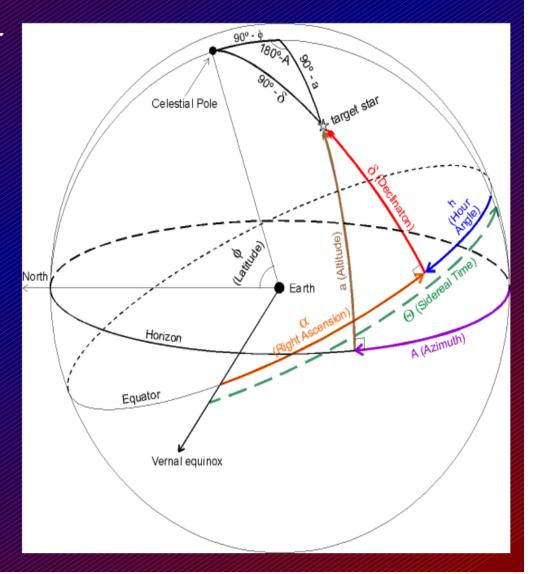


Sidereal Time and Hour Angle

Sidereal Time = RA of stars crossing the meridian or HA of Vernal Equinox.

For any star: ST = RA+HA

Sidereal Day = time between two successive meridian crossings of VE.



Spherical Trigonometry

To convert from equatorial to horizontal coordinates:

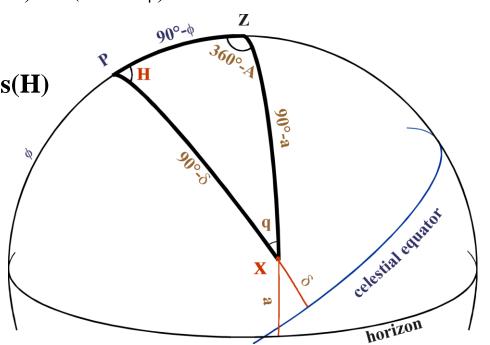
Given RA α and declination δ , we have Local Hour Angle H = LST - RA, in hours; convert H to degrees (multiply by 15). Given H and δ , we require azimuth A and altitude a. By the cosine rule: $\cos(90-a) = \cos(90-\delta) \cos(90-\phi) + \sin(90-\delta) \sin(90^{\circ})$ -0) $\cos(H)$ Ζ 90°-¢ which simplifies to: $sin(a) = sin(\delta) sin(\phi) + cos(\delta) cos(\phi) cos(H)$ This gives us the altitude a.

By the sine rule:

```
sin(360-A)/sin(90-\delta) = sin(H)/sin(90-a)
which simplifies to:
```

```
-\sin(A)/\cos(\delta) = \sin(H)/\cos(a)
```

```
i.e. \sin(\mathbf{A}) = -\sin(\mathbf{H})\cos(\delta) / \cos(a)
which gives us the azimuth A.
```



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Make changes then press Update to run.				
Local		MST Calendar		
Tucson, Arizona		August 2018		
Latitude:	32:13:15	Su No Tu He Th Fr Sa		
Longitude:	110:58:08	29 30 31 1 2 3 4		
Elevation:	728,5 n	5 6 7 8 9 10 NH 12 13 14 15 16 17 18		
Temp:	10.0 C	19 20 21 22 23 24 25		
Atn Pres:	1010 hPa	FH 27 28 29 30 31 1		
Equinox:	2000.0	2 3 4 5 6 7 8		
Mag decl:	0:00:00	<< < Nov > >>		
Time 🕐 🔿 Night				
Julian:	2458357,55676	Sun Dip: 18°		
UTC Date:	8/27/2018	Dawn: 4:29		
UTC Time:	1:21:44	Dusk: 20:21		
Sidereal:	16:18:48	Length: 8:08		
TZ Nane:	MST	LST00: 21:58:00		
TZ Offset:	7:00:00	-1+1 Looping RT		
Local Date:	8/26/2018	Step: Clock		
Local Time:	18:21:44	N Steps: 1		
Delta T:	(Auto) 69,85	Pause: 0		

Let's practice at an actual telescope: Our 21"

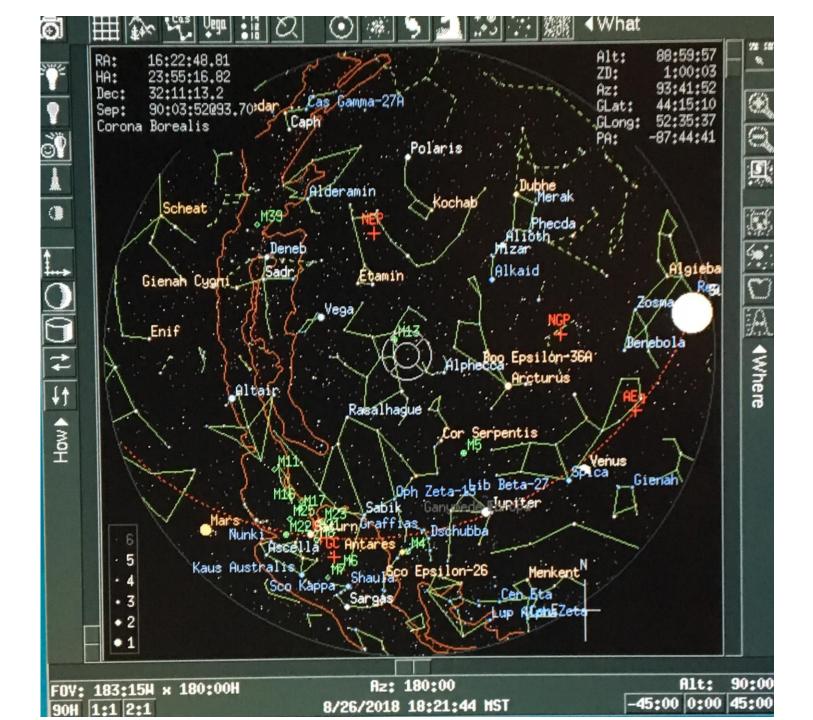
Steward's Telescope Control System (TCS)

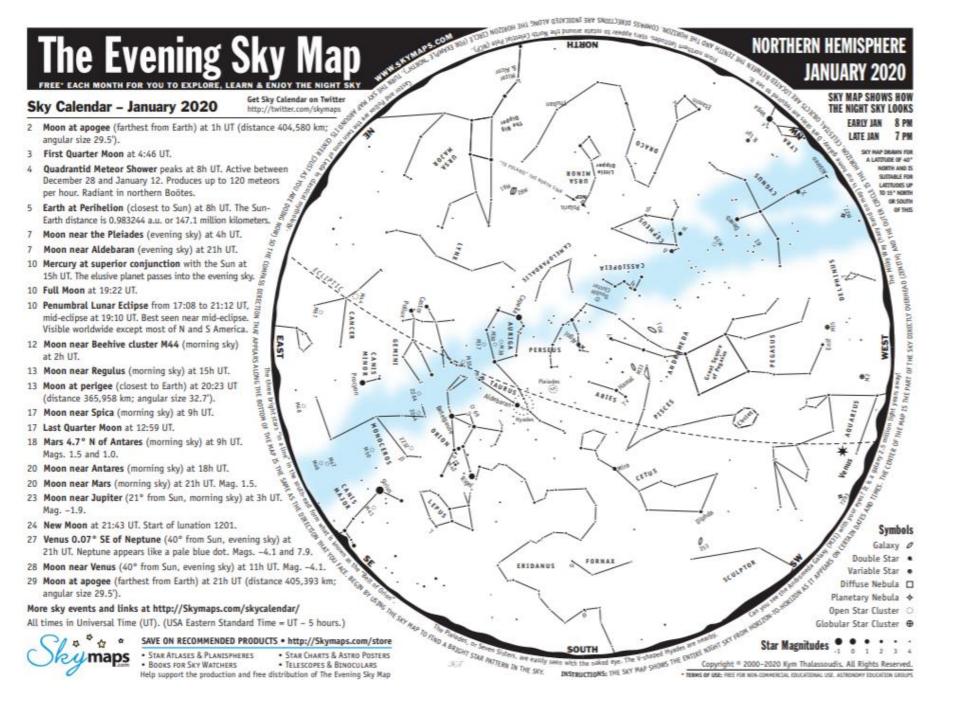
Erase Trace to indit	race.txt	
TCS-NG-INDI		
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Hour Angle		+00:00:1
Sidereal Time		16:23:0 89.
Elevation		85. 180.
Azinuth		1.0
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Universal Time		01:25:52.8
Corrections		MPNARFp+
Axial Telenetry		
	RA	DEC
	HH:MM:SS.ss	D:HH:SS.
Current	16:22:05.25	
Connanded	16:22:05.26	+32:11:14.
Next	16:17:16.30	+32:12:01.
Offset difference	+00:00:00.01	+00:00:00.0
ditterence	-0.015 RA	-0,00 TFS
bias	0.000	0.00
guide	40.000	40.000
drift	1000.000	1000.000
Actions		
Go Next Cancel Enable		
Zero Tracking Offset		
🖽 Goto Functions		
Corrections		

[7] n .

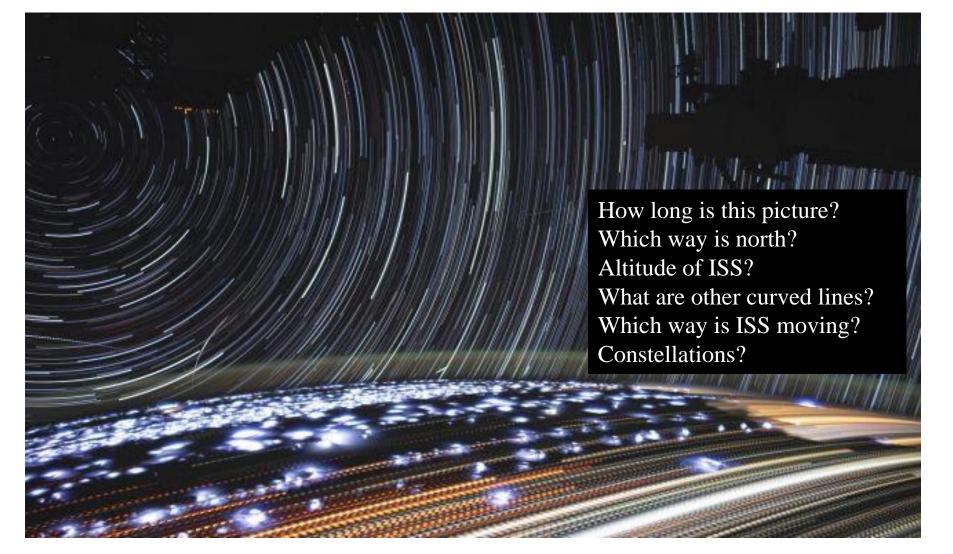
Where is the telescope pointed now?

Small angular corrections are applied





Where was this picture taken? Exposure time = ?



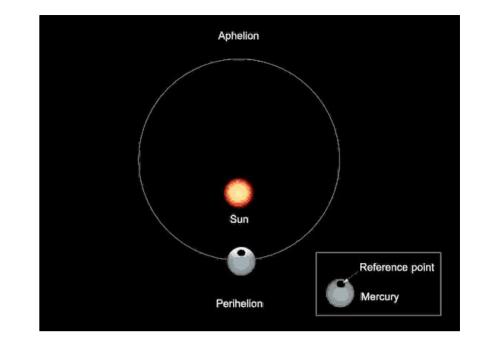
Mercury: Solar vs. Sidereal Day

How long is a "day" on Mercury?

Mercury rotates on its axis exactly 1.5 times for each revolution around the Sun.

> Revolution period 87.969 days

Rotation period 58.646 days



Mercury rotates exactly 3 times during two orbits around Sun.

RATIO = 87.97/58.65 = 3:2 ["3:2 spin-orbit resonance"]

What does it mean to "revolve"? How long is a "year"?

- Synodic period = 29.53 days
- Sidereal period = 27.32 days

